

Ch 11 - Kinetics

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First, take $m_e \rightarrow 0^+$ in the electron eqn of motion (i.e., electron is massless, no inertia), we have

$$\underline{E} + \underline{v} \times \underline{B}/c = 0$$

This implies $\underline{B} \cdot \underline{E} = 0$; i.e. electrons are infinitely mobile along \underline{B} to short circuit any electric field parallel to \underline{B} . Summing up the eqn. of motions of both ions and electrons, we find

$$\rho_i \frac{d\underline{v}}{dt} = \underline{J} \times \underline{B}/c ;$$

where we have applied quasi-neutrality $n_e = n_i$ to eliminate the electric field, and noted again the plasma is taken to be "cold". $\rho_i = n_i m_i$ and inertia is carried by ions. Quasi-neutrality also implies

$$\underline{\nabla} \cdot \underline{J} = 0.$$

Consider infinitesimally small (linear) perturbations from the following stationary equilibrium

$$\begin{aligned} \underline{v}_0 = 0 &\Rightarrow \underline{v} = \delta \underline{v} \\ \underline{E}_0 = 0 &\Rightarrow \underline{E} = \delta \underline{E} \\ \underline{J}_0 = 0 &\Rightarrow \underline{J} = \delta \underline{J} \\ \underline{B}_0 \neq 0 &\Rightarrow \underline{B} = \underline{B}_0 + \delta \underline{B}, \quad n_0 \neq 0. \end{aligned}$$

δ 's are 0^+
linear quantities

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Thus, $\delta \underline{V} \times \underline{B}_0 / c = -\delta \underline{E} \Rightarrow \delta \underline{V}_\perp = c \delta \underline{E} \times \underline{B}_0 / B_0^2$.

$$\rho_{i0} \frac{\partial}{\partial t} \delta \underline{V} = \delta \underline{J} \times \underline{B}_0 / c \Rightarrow \delta \underline{J}_\perp = - \frac{\rho_{i0} c}{B_0^2} (\partial_t \delta \underline{V}_\perp) \times \underline{B}_0$$

Here, \perp denotes (perpendicular to \underline{B}_0) component.

$$\therefore \boxed{\delta \underline{J}_\perp = \frac{\rho_{i0} c^2}{B_0^2} \frac{\partial}{\partial t} \delta \underline{E}_\perp}$$

$$\boxed{\underline{\nabla} \cdot \delta \underline{J} = 0} \quad \text{with} \quad \delta \underline{J} = \frac{\delta J_{\parallel}}{B_0} \underline{B}_0 + \delta \underline{J}_\perp \quad \parallel = \text{parallel to } \underline{B}_0$$

$$\Rightarrow \boxed{\underline{B}_0 \cdot \underline{\nabla} \frac{\delta J_{\parallel}}{B_0} + \underline{\nabla}_\perp \cdot \delta \underline{J}_\perp = 0}$$

Adopt scalar and vector potentials,

$(\delta \phi, \delta \underline{A})$ Coulomb gauge $\underline{\nabla} \cdot \delta \underline{A} = 0$.

$$\delta \underline{E} = -\underline{\nabla} \delta \phi - \frac{1}{c} \partial_t \delta \underline{A}$$

Drupere's Law

$$\nabla^2 \delta A_{\parallel} = -4\pi \delta J_{\parallel} / c$$

Alfvén waves have phase velocity \ll speed of light \Rightarrow displacement current is negligible.

Furthermore, we assume

$$|k_{\parallel}| \ll |k_{\perp}| \quad \text{or} \quad \left| \frac{\underline{B}_0}{B} \cdot \underline{\nabla} \right| \ll |\underline{\nabla}_\perp|;$$

again in the spirit of minimize $|k_{\parallel}|$.

$$\Rightarrow \nabla^2 \approx \nabla_\perp^2, \quad \text{and} \quad \delta \underline{E}_\perp \approx -\underline{\nabla}_\perp \delta \phi$$

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$$\underline{\nabla} \cdot \underline{\delta J} = 0 \Rightarrow$$

$$c \underline{B}_0 \cdot \underline{\nabla} \frac{\underline{\nabla}_\perp^2 \delta A_\parallel}{B_0} - 4\pi \underline{\nabla}_\perp \cdot \underline{\delta J}_\perp = 0$$

$$\delta E_{\parallel} = 0 = \underline{B}_0 \cdot \underline{\nabla} \delta \phi \Rightarrow \boxed{\underline{B}_0 \cdot \underline{\nabla} \delta \phi = \frac{1}{c} \partial_t \underline{B}_0 \cdot \underline{\delta A} = \frac{B_0}{c} \partial_t \delta A_{\parallel}}$$

Denote: $\underline{B}_0 \cdot \underline{\nabla} = B_0 \frac{\partial}{\partial \ell}$

$$\Rightarrow \boxed{B_0 \frac{\partial}{\partial \ell} \frac{1}{B_0} \nabla_\perp^2 \frac{\partial}{\partial \ell} \delta \phi - \frac{1}{v_A^2} \frac{\partial^2}{\partial t^2} \nabla_\perp^2 \delta \phi = 0}$$

with $v_A^2 = \frac{B_0^2(r, \theta)}{4\pi \rho_0(r)} = \frac{B_0^2}{4\pi \rho_0(r)} \cdot \left(1 + \frac{r}{R_0} \cos \theta\right)^{-2}$.

⊙ It thus suggests the following general wave equation

$$\boxed{\mathcal{L}(\omega, \frac{\partial}{\partial \ell}, \nabla_\perp^2, \cos \theta, r) \delta \phi(r, \theta, \xi) = 0,}$$

where we have let $\frac{\partial}{\partial t} = -i\omega$, etc.

Taking, with the translational invariance constraint,

$$\delta \phi(r, \theta, \xi) = e^{in\xi} \sum_m e^{-im\theta} \delta \Phi(n\xi - m)$$

$$\Rightarrow \text{we have } e^{in\xi} \sum_m e^{-im\theta} \mathcal{L}(\omega, ik_{\parallel nm}, \frac{\partial^2}{\partial r^2} - \frac{n^2 g^2}{r^2}, T_c, r_m) \delta \Phi(n\xi - m)$$

with $T_c \delta \Phi(n\xi - m) = \frac{1}{2} [\delta \Phi(n\xi - m - 1) + \delta \Phi(n\xi - m + 1)]$.

p. 16 With trans. inv., the linear wave eqn becomes

$$\mathcal{L}(\omega, i k_{||nm}, \partial_r^2 - \frac{n^2 g^2}{r_m^2}, T_c, r_m) \delta \bar{\Phi}(n\theta - m) = 0.$$

$k_{||nm} = \frac{1}{gR}(n\theta - m)$. Note it is a differential-difference equation. Difference operator is due to T_c .

This equation becomes more readily solvable if we Fourier transform $\delta \bar{\Phi}(n\theta - m)$. !!

• Defining $\delta \hat{\Phi}(y) = \text{F.T. of } \delta \bar{\Phi}(n\theta - m)$; i.e.

$$\delta \hat{\Phi}(y) = \int \delta \bar{\Phi}(n\theta - m) e^{i(n\theta - m)y} d(n\theta - m)$$

$$\delta \bar{\Phi}(n\theta - m) = \frac{1}{2\pi} \int \delta \hat{\Phi}(y) e^{-iy(n\theta - m)} dy$$

$$\Rightarrow i k_{||nm} \delta \bar{\Phi}(n\theta - m) = \frac{1}{gR} i(n\theta - m) \delta \bar{\Phi}(n\theta - m) \xrightarrow{\text{F.T.}} \frac{1}{gR} \frac{\partial}{\partial y} \delta \hat{\Phi}(y)$$

$$\Rightarrow (\partial_r^2 - \frac{n^2 g^2}{r_m^2}) \delta \bar{\Phi} \xrightarrow{\text{F.T.}} \left[-\frac{n^2 g^2}{r_m^2} + (-iy n g')^2 \right] \delta \hat{\Phi}(y)$$

$$= -\frac{n^2 g^2}{r_m^2} [1 + \hat{\zeta}^2 y^2] \delta \hat{\Phi}(y)$$

$$\Rightarrow T_c \delta \bar{\Phi}(n\theta - m) \xrightarrow{\text{F.T.}} (\cos y) \delta \hat{\Phi}(y) \quad \hat{\zeta} = \frac{r_m g'}{g} \text{ magnetic shear}$$

$$\Rightarrow \mathcal{L}(\omega, \partial_y, -\frac{n^2 g^2}{r_m^2} (1 + \hat{\zeta}^2 y^2), \cos y, r_m) \delta \hat{\Phi}(y) = 0 \quad !!$$

ordinary differential eqn. parameterized by r_m .

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We need to impose appropriate boundary conditions to uniquely specify the above linear eigenmode equation.

One of the most useful consideration is taking the initial-value view point; i.e., the causality constraint.

That is, $\delta\phi(t) \rightarrow 0$ as $t \rightarrow -\infty$. Taking $\delta\phi(t) \propto \hat{\delta\phi}(\omega) \exp(-i\omega t)$, we then dictate that

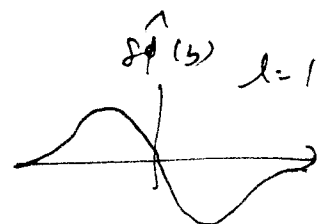
$$\hat{\delta\phi}(\omega) \rightarrow 0 \text{ as } |\omega| \rightarrow \infty \text{ for } \text{Im}\omega > 0.$$

Physically, one can understand this causality boundary condition by noting that, since the group velocity is finite, any initial perturbations ^(t=0) with finite extent in y will not reach $|y| \rightarrow \infty$ during the finite t . Imposing the b.c.'s, we can, in principle, solve the eigenvalue and eigenfunction as

$$\omega = \omega(r_m | n, l) \quad l = 0, 1, 2, \dots$$

$$\hat{\delta\phi} = \hat{\delta\phi}(y | n, l, r_m) \quad ;$$

here l is the l th eigenstate; e.g.



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o For the shear Alfvén wave discussed earlier,
the eigenmode equation is

$$\left\{ \frac{\partial}{\partial y} k_{\perp}^2 \frac{\partial}{\partial y} + \frac{\omega^2}{\omega_A^2(r_m)} k_{\perp}^2 \left[1 + \underbrace{\left(\frac{r_m}{R_0} \cos(y) \right)}_{\textcircled{4}} \right] \right\} \delta\phi(y) = 0;$$

where $k_{\perp}^2 = \frac{n^2 g^2(r_m)}{r_m^2} [1 + s^2(r_m) y^2]$, $\omega_A^2 = \frac{V_A^2(r_m)}{g^2(r_m) R_0^2}$.